# RAMAN RESEARCH INSTITUTE BANGALORE 560 080

# ANNUAL REPORT 1989-90

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## Introduction

The Raman Research Institute founded by Prof. C.V. Raman in the late forties was reorganised, after his death in 1970, as a national Institute for research in basic science and it has been receiving generous grants from the Department of Science and Technology of the Government of India since 1972. The main fields of research are Astronomy and Astrophysics, and Liquid Crystals.

# 1. Astronomy and Astrophysics

Astronomy which is one of the oldest sciences is concerned with the study of heavenly bodies by investigating the radiation received on earth from them. Optical Astronomy deals with the "visible" part (wavelength 3000 Angstroms to 6500 Angstroms, 1 Angstrom =  $10^{-8}$  centimeter) of the electromagnetic spectrum. Radio Astornomy, which had its beginnings in 1932 also deals with the study of these heavenly bodies, but the radiation received by radio telescopes on earth is in the radio wavelength part (30 meters to 1 millimeter) of the same electromagnetic spectrum. The lower and upper limits in wavelength of the radio spectrum are set by the earth's atmosphere and ionosphere respectively. In spite of these limitations, the radio window is very wide (30,000 to 1 compared to the 2:1 wavelength ratio in the visible part) and studies within it over the years have yielded information leading to many exciting discoveries such as the 3° k cosmic background radiation, quasars, pulsars, etc.

The Raman Research Institute has observational programmes in Radio Astronomy extending over most of the available radio spectrum. It has setup a Decameterwave Radio Telescope at Gauribidanur jointly with the Indian Institute of Astrophysics. Operating at a wavelength of 10 meters, it is among the largest telescopes in the world and is being used to study the radio emission from various types of celestial objects such as the Sun, Jupiter, radio sources of various kinds in our Galaxy, and external Galaxies. Moving to somewhat shorter wavelengths, members of the Institute use the Ooty Radio Telescope operated by the Tata Institute of Fundamental Research, Bombay. This instrument operates at a wavelength of approximately 1 meter and is used for carrying out observations of pulsars, and nebulae of various kinds in the Galaxy.

During the past two decades, millimeterwave astronomy has assumed great importance because of the discovery of numerous molecules (combinations of Hydrogen,

Carbon, Nitrogen, Oxygen, Silicon, etc.) by their emitted line radiations in the shortest wavelength region of the radio spectrum. These molecules are generally found in dense molecular clouds in our own and other Galaxies where star formation is thought to be taking place. The Raman Research Institute has setup two millimeterwave telescopes of diameters 1.5 and 10.4 meters, which are used for the study of these radiations.

In addition, the Institute has theoretical research programmes in many branches of Astrophysics like Pulsars, Supernova Remnants, the interstellar medium, Galaxies and several aspects of General Relativity and Gravitation.

# 2. Liquid Crystals

Liquid Cystals are states of matter intermediate between the liquid and cystalline states. Many organic compounds whose molecules have pronounced shape anisotropy exhibit such phases. The unique combination of fluidity and anisotropic properties of liquid crystals has led to many applications of these materials. The Liquid Crystals Laboratory of the Raman Research Institute has contributed significantly to the development of the field over the past decade and a half.

The laboratory has been organised to undertake studies of most of the fundamental properties of liquid crystals. Theoretical and experimental work on liquid crystals is continuing along the lines indicated in the reports of previous years.

A somewhat more technical account of the work carried out in the past year is given in the following.

#### THEORETICAL PHYSICS

Majorana Fermions: It is known that in the four dimensional Euclidean space there are no Majorana fermions. A generalisation of the Majorana condition was found which enables one to define "Majorana-like fermions" in Euclidean space. It turns out that an extended supersymmetric Yang-Mills (SSYM) theory constructed out of these fermions becomes an ordinary SSYM or a topological SSYM depending on the choice of the definition of the Majorana-like fermions.

Maximum Entropy Method: This technique has recently attracted some interest in the Crystallographic Phase Problems. In this context MEM is viewed as a variational way of obtaining a density map which agrees simultaneously with experimental data, as well as with physical constraints based on a priori knowledge of the characteristics of the density function. For two specific "entropy" functions, the topology of the "entropy surface" in the space of unknown phases was investigated numerically. Multiple extrema were found to be the rule. A hybrid algorithm based on the gradient search and incorporating the simulated annealing philosophy was tried on a heavy atom real structure and showed promising results. The implications of these results for potential numerical algorithms is being explored.

Geometric Phase and Optics: The problem of synthesizing general polarization transformers with quarterwave and halfwave plates was pursued further and a five-element combination with several useful properties discovered. For example, it allows the practical realization of an exact analogue of an arbitrary hamiltonian acting on a spin 1/2 system. Using this device, phase changes arising in the evolution of spin 1/2 systems under several different kinds of hamiltonian were simulated in a HP laser interferometer system, thus demonstrating the concept of a 'spin 1/2 interferometer' using light beams. Further applications of this concept are being investigated.

It was found that the method of decomposing arbitrary rotations into desirable ones used in the synthesis of polarization retarders can be applied to the problem of 'zero angular momentum turns' involved in phenomena like the turning of cats spacemen and divers, yielding useful insight.

### GENERAL RELATIVITY AND COSMOLOGY

Geodesic motion in higher dimensional Boulware-Deser black hole spacetime was studied and the results compared with those for higher dimensional Schwaarzschild geometry. It was found that for both time-like and null geodesics no suitable bound orbit is possible.

The properties of black hole solutions in higher dimensions are of importance in the study of string theories and the quantum theory of gravity. Motivated by this, a study of scalar perturbations of a Kerr-type black hole in higher dimensions has been initiated.

Previous work on approximating the black hole potential by a patchwork of potentials for which the scattering problem may be exactly solved was continued. The effective potential for the Schwarzschild black hole was approximated by two Eckart potentials and a  $r^{-2}$  potential. The solution was obtained in terms of hypergeometric and confluent hypergeometeric functions. Matching the solutions yielded an analytic expression for the reflection coefficient. The results thus obtained agreed well with numerical results.

The Relativistic multipole expansion for electromagnetism and linearized gravity in the region outside a spatially compact supported source has been obtained directly within the STF formalism. The results were compared with those obtained earlier using the Debye potential formalism. The slow motion limit of the mass and spin moments agree with those of Thorne at the first Post Newtonian level.

Gravitational Radiation: Starting from the first Post Newtonian expressions for radiative spin moments obtained in linearised gravity, the effects of quadratic non-linear terms were taken into account. A prescription to obtain compact support expressions for the moments was introduced and used to study the spin moments. This completes the discussion of the emission of gravitational waves upto order  $(v/c)^3$  beyond the Einstein-Landau-Lifshitz quadrupole formula.

Massless scalar waves in the Witten Bubble spacetime have been investigated. The time like and angular parts of the separated Klein Gordon equation are written in terms of hyperbolic harmonics characterised by a generalised frequency. The radial equation is cast into the Schrodinger form. The above mathematical formulation has been applied to study the scattering problem, the bound states and the corresponding stability criteria. The results confirm the concept of a bubble wall as a perfectly reflecting expanding sphere.

### LIQUID CRYSTALS

Work is continuing as usual on diverse aspects of the multidisciplinary field of liquid crystals. A brief summary of some of the most significant results is given below.

It was mentioned in last year's report that a notable achievement of the laboratory was the prediction and discovery of a biaxial nematic liquid crystal phase in a simple thermotropic system. Since then two other cases of the biaxial nematic phase have been reported, both from Germany. In all three cases the molecule combines the features of the rod and the disc as was predicted here some years ago. Recently, it has been possible to obtain further confirmatory evidence of the biaxial nature of the nematic phase of our compound by X-ray methods. It is well known that the classical uniaxial nematic phase gives rise to two liquid-like X-ray diffraction maxima. In the present case the diffraction pattern consists of there diffuse maxima as would be expected from an orthorhombic fluid.

Theoretical investigations on electrohydrodynamic instabilities in nematic liquid crystals were continued. It has been shown that if the symmetry of the cell is changed by incorporating a small tilt angle of the director at the boundary surfaces, the instability pattern propagates. This mode has been observed by us experimentally.

A simple linearized perturbation theory has been developed to account for electric field induced static modulated structures in nematics. Theoretical calculations are in qualitative agreement with experiment with regard to the threshold, domain size, etc. A rigorous numerical approach has been adopted to study the occurrence of the "stripe phase" above the bend Freedericksz transition in nematics at temperatures close to the smectic A transition. While qualitatively accounting for the experimental observations regarding the domain of existence of the stripe phase, the calculations also indicate the possibility of the occurrence of stripes in certain polymer nematic systems such as the tobacco mosaic virus (TMV).

The influence of a magnetic field on the structure and properties of defects in liquid crystals has been studied theoretically. Interesting new results are predicted in certain situations. For example, a pair of parallel disclination lines in nematics subjected to a magnetic field interacts through a distance-independent force law, which is quite different from the 1/r law of interaction for the field free case. In diamagnetically negative nematics, at high fields, a singular line of strength S=+1 terminates within the body of the material. In the case of smectics, in the field free case, screw dislocations have neither self energy nor do they interact. However, when a field is applied they interact through a 1/r force law; also smectics can exhibit

screw dislocation mediated instabilities. In the case of discotic liquid crystals a S=+1 defect with the columns encircling the disclination line exhibits a structural instability in a magnetic field.

The director configuration in the nematic-nematic co-existence range of mixtures of rod-like and disc-like compounds was investigated. Using these studies, the topological equivalence of point defects of strength +1 and ring defects of strength +1/2 could be demonstrated experimentally.

About forty new ferroelectric liquid crystals have been synthesized. Some of these have fairly low melting points and moderately high value of spontaneous polarisation. A ferroelectric liquid crystal containing a transition metal has also been synthesized. The properties of this class of materials are under investigation.

Although in recent years ferroelectrics have been attracting a great deal of attention, the relationship between molecular structure and physical properties is still not clearly understood. In particular, information about the influence of molecular structure on spontaneous polarisation, response time and rotational viscosity are important from the applications point of view also. In order to study these properties in detail a completely computerised high resolution set up has been constructed. Using this set up experiemnts have been conducted on four different but structurally related homologous series of compounds synthesised by this laboratory. The results have led to a better understanding of the structure – property relationship.

Several new metal complexes have been prepared and some of them exhibit interesting polymesomorphism. Uniaxial nematic, biaxial nematic, smectic A, smectic C, smectic C\* and smectic E phases have been identified in some of these complexes.

The crystal and molecular structure of a discogen, bis[1,3-di(p-n-decyl-phenyl)propane-1,3-dionato] palladium (II) has been analysed. The sample crystallizes in the triclinic space group  $^-P1$  with one molecule in a unit cell. The crystal structure was solved by direct methods and refined using full matrix least squares procedure. The palladium atom is located at an inversion centre, and the decyl chains are fully extended in an all trans conformation. In the crystal structure, the molecules form a tilted columnar arrangement, each column being surrounded by six others. The crystal molecular structures of other organometallic complexes which form discotic liquid crystals are also being investigated.

In the area of applications, the possibility of lowering the supply voltage requirement without sacrificing the contrast ratio of multiplexed displays by making use of generalised addressing scheme, is being explored.

The Liquid Crystal Laaboratory, Raman Research Institute, was recently conferred a unique honour on the occasion of the 900th anniversary celebrations of the University of Bologna, Italy. The laboratory was "awarded a special prize in a world-wide selection of the most distinguished Academic and Scientific Institutions in the world". Professor Chandrasekhar was invited to Bologna to receive the award at a special ceremony held on 9 June 1989. The Laboratory shared this honour with about a

dozen other famous centres which included CERN (Geneva), ICTP (Trieste), the Institut Pasteur (Paris), the Laboratory for Molecular Biology (Cambridge), etc., and about 25 of the world's most renowned Universities.

#### ASTRONOMY AND ASTROPHYSICS

Stellar Dynamics: The study of the dynamics of oscillating three dimensional collisionless stellar systems (shells, ellipsoids) was continued.

The dynamics of planar systems which are time dependent generalisations of Freeman discs was also investigated. It was found that in the tidal approximation, the distribution function contains ten parameters whose evolution specifies a system of ordinary differential equations. Results were obtained for energy and angular momentum transfer.

Pulsars: A novel mechanism was suggested for the underlying physical mechanism responsible for the decay of the magnetic fields of neutron stars. This exploits the fact that both the superconducting protons and the suerfluid neutrons in the interior of a magnetised rotating neutron star will be in a vortex state. The secular decay of the magnetic field is related to the secular slow down of the neutron star, through a strong interpinning between the vortices in the superconductor and the superfluid. Among many other things, this model is able to explain for the first time why low field neutron stars are almost always in binary systems.

It was pointed out that the fact that the measured mass of the millisecond pulsar PSR 0021-72A in the globular cluster 47 Tucanae is close to the Chandrasekhar mass lends strong support to the possibility that the radiation from pulsars spun up to ultra short periods may be efficient in evaporating the companion. If this is true, then the lifetime of low mass X-ray binaries may be 10 to 100 times smaller than generally believed, alleviating many evolutionary problems concerning the progenitors of millisecond pulsars.

One of the controversial questions in the literature has been whether the spin axis and magnetic axis of pulsars align with age. Using the recently available data on the inclination angles of nearly 150 pulsars it was demonstrated that there is no observational evidence to support the alignment hypothesis.

Effect of limited data length on the estimates of slow intensity fluctuations resulting from refractive effects in the interstellar medium was studied. The results of Monte Carlo simulations indicate that the apparent modulation index reduces if the duration of the observations is not very long compared to the refractive time scale. These computations can be used to estimate the true modulation index given the measured values.

A new observational programme for detailed studies of pulsar radiation at meter wave length using the Ooty radio telescope has been started in collaboration with the TIFR group. The development of appropriate software to enable the acquisition, and on line processing of pulsar data, has been completed. Many important modifications

were made to an available digital correlation receiver to suit pulsar observations, and the new receiver system has been successfully tested.

The design and construction of suitable instrumentation for pulsar searches and observations with the Giant Meter Wavelength Telescope (being built near Pune) is in progress.

Radio Emission from the Galactic Centre: Using the Very Large Array radio telescope, a  $2^{\circ} \times 2^{\circ}$  wide field around the galactic centre was imaged at 332 MHz with angular resolution ranging from 15'' to 100''. The images show all the unique features associated with the Galactic center, including the linear filaments, and several long "thread-like" features. Many linear filaments in this region, although non-thermal, show postive spectral index ( $\infty = +0.3$ ). This may indicate in situ production of relativistic electrons on time scales of less than  $10^4$  years. The orientation of the "threads" and the filaments suggest the presence of a strong poloidal magnetic field in the central 100 pc of the Galaxy.

Carbon recombination lines towards Cas A: These were investigated over the frequency range 34-325 MHz, corresponding to transitions in the quantum number range n=578 to n=272. The lines appear in absorption at low frequencies (<150 MHz) and in emission at high frequencies (>200 MHz) as expected. In order to determine the origin of these lines, optical depth maps of  $C272 \approx (325 \text{ MHz})$  were obtained using the VLA with a resolution of  $\sim 1.5'$ , and compared with those of 21 cm HI and 6 cm  $H_2CO$  absorptions. The recombination line maps are similar to 21 cm HI absorption maps suggesting that the lines originate in warm (50-100K) HI clouds with an electron density  $\sim 0.05 \text{ cm}^{-3}$ . However, the variation of the intenstiy of the line with frequency is better fit by a model in which the lines originate in cold (15-20K) regions with electron densities  $\sim 0.3 \text{ cm}^{-3}$ . This dilemma is yet to be resolved and further observations near 600 MHz are planned.

The warm ionized interstellar medium: A hydrogen recombination line near 145 MHz was detected in the direction of the galactic centre. The strength of this line not only clarifies the nature of a nearby HII region, but also puts strong constraints on the properties (such as the temperature and the filling factor) of the diffuse warm ionized gas.

Recombination lines from the Star Burst Galaxy NGC 253: NGC253 was observed in  $H166 \infty$ ,  $H110 \infty$  and  $H92 \infty$  recombination lines using the VLA. Lines were detected in the neclear region, which also dominates the continuum emission. The location of the  $H166 \infty$  line emitting region is different from that of the other two. It is suggested that the higher frequency lines arise in high density and high emission measure gas, and the lower frequency line in a region of low density and lower emission measure. The observed velocity distribution indicates unusual motions of the ionized gas near the nucleus.

Orion B: The spatial and velocity distribution of the narrow  $H166 \propto$  and  $C166 \propto$  lines in Orion B, observed with an angular resolution of  $\sim 45$ ", were found to be similar.

This strongly suggests that the CII and  $H^{\circ}$  regions, which are adjacent to the HII region, are spatially coincident. The heavy element line (X166 $\infty$ ) has a different distribution and its identification as due to sulfur, based on its velocity, is not certain. A new H166 $\infty$  feature was detected at a velocity of  $+48kms^{-1}$  in the south-western portion of Orion B.

# Millimeterwave Astronomy

# 10.4 meter telescope

Observations: Very Long Baseline Interferometric (VLBI) observations of strong water vapour maser sources at 22 GHz have been carried out using our 10.4m telescope, the 22m telescope in Crimea, USSR and the 100m telescope in Bonn, West Germany – during September, 1989. These observations were only intended for testing the RRI system and the problem areas were identified and rectified. Main observations are planned for May, 1990.

Further observations on Mira variables by SiO maser emission in the  $V=1,\,J=2\to 1$  transition at 86 GHz were continued during the first half of the observing session which had to be cut short by bad weather in the initial period. During this session 50 more sources were observed. Besides supporting the earlier conclusions about the dependence of maser emission on the spectral type of the star, it was found that the maser luminosity varies inversely with the amplitude of variation of the visual magnitude. Interpretation of these results is in progress.

<sup>12</sup>CO and <sup>13</sup>CO observations were also undertaken during the last three weeks of the observing session on Herbig Haro objects and Lynds' clouds. Under a collaborative programme with the Infrared Astronomy Group of the Tata Institute of Fundamental Research, Bombay, two regions W31 and G351.6-1.3 were selected for mapping in <sup>12</sup>CO. Limited observations indicated that the spectra are complex and show a wealth of information. It is hoped to undertake detailed mapping during the next observing session.

Another observational project, using the 10.4m telescope in conjunction with the newly developed 400 MHz bandwidth acousto-optic spectrometer, was the search at 104 GHz, for  $SO_2$  molecules in the atmosphere of Venus. New calibration techniques for observation of broad lines in the presence of strong continuum radiation were tested. The data is under reduction.

Receiver and Component Development: A dual polarisation cooled receiver operating in the band 86 to 115 GHz was fabricated and tested in the laboratory. Initial tests on the telescope revealed several minor problems which have been solved. Further tests will be carried out on the telescope. Four more GaAs FET amplifiers operating at the intermediate frequency of 1.45 GHz with an instantaneous bandwidth of 400 MHz have been fabricated, tested and optimised for cryogenic operation at an ambient temperature of 20K. Two more Schottky barrier diode mixer blocks for the frequency range 86 to 115 GHz have also been fabricated and optimised for cryogenically cooled operation. A tunable Gunn oscillator for the frequency range 80 to 114 GHz has also been fabricated and tested. It yields enough power for operation of cryogenically cooled mixers even at the higher frequency end which is the difficult region.

A cooled 22 GHz receiver for observations of galactic water vapour maser sources has been assembled using a circular polariser ahead of a HEMT amplifier obtained on loan from the National Radio Astronomy Observatory. USA. In order to use this receiver at the nasmyth focus of the 10.4 metre radio telescope, a 2.3 metre long conical horn with an aperture dia of 26 cm was designed and fabricated. The system was primarily intended for observations in the Very Long Baseline Interferometric (VLBI) mode between Crimea in the USSR, Bonn in West Germany and Bangalore. For these observations, two Hydrogen Maser frequency standards and a Mark II VLBI data recording system have been supplied by the Space Research Institute, Moscow, USSR.

A sensitive CCD camera system was designed based on a Philips camera module of 0.05 lux sensitivity and was used along with a Questar telescope for optical pointing of the 10.4 m telescope. A pointing correction model was obtained using this system by observing stars upto the sixth magnitude.

Data acquisition: The computer which controls the 10.4 metre telescope was enhanced by incorporating a faster CPU system and additional memory so that both the control operations and preprocessing of the on line data could be performed in real time. The software for controlling the telescope and acquiring data from the receiver was suitably modified.

A microcomputer based on Motorola 68000 processor was developed and was used in the current observing session to act as slave computer to acquire data from both the 64 MHz filter bank and 40 MHz Acousto-optic spectrometer (AOS). Stability tests of all the back-end spectrometers were carried out with an Allan variance test procedure.

#### Decameterwave Astronomy

All sky map at 34.5 MHz: Further analysis was done on the recently completed sky survey at 34.5 MHz. The survey was found to be confusion limited at 5 Jy. The maps which have a resolution of  $26' \times 42'$  at zenith, were smoothed to a resolution of  $2^{\circ}$  to highlight the large angular scale features. Several well known Galactic loops and spurs were identified in the 34.5 MHz maps and their spectra were obtained by comparing with already existing 408 MHz observations. All the spectra indicated the non-thermal nature of the loops and spurs. The well known warp of the galactic plane was also clearly seen in the 34.5 MHz map.

Absorption of the galactic background emission due to the foreground thermal gas in the Galaxy was seen clearly in the inner Galaxy. There was no sign of such absorption in the direction of the outer Galaxy. Several discrete absorption regions were detected in the present survey. The nature of the gas causing these absorptions was studied. There is strong evidence to suggest that this gas is fairly nearby ( $\leq 3$  kpc), extended (20-100 pc) and of low density (10 electrons  $cm^{-3}$ ). Most of the low velocity ( $\leq 50$  kms) H166 $\propto$  and H272 $\propto$  galactic ridge recombination lines can be easily accounted for as arising in this nearby gas. This interpretation has thrown considerable doubt on the core-envelope picture of HII regions.

Solar Observations: The study of type III solar bursts using the broad band array in Gauribidanur (operating at the range 35 MHz to 65 MHz) is being continued. The frequency drifts with time of Soalr bursts are being studied using a 1760-channel acousto-optic spectrometer. Such studies will help in understanding the movement of charged particles in the Solar corona.

Long baseline interferometry: The construction of two receiver systems for acquiring data from two stations has been completed. Data will be recorded at each of the two stations, and the correlation will be performed off-line. Initial tests with a 5 kilometer baseline have yielded satisfactory results.

#### MAURITIUS TELESCOPE PROJECT

An interferometer using 32 helices with a baseline of 700 meters, operating at a frequency of 150 MHz is operational at Gauribidanur. After finalising the sub-assemblies used at this prototype system, equipment necessary for a similar set-up at Mauritius has been fabricated. Installation of this system has just started there.

Part of the equipment from the Clark Lake Radio Observatory (CLRO), which includes the Front-end post amplifiers and Two-stage mixers has been fully tested and reconfigured to suit the requirement of the Mauritius synthesis array.

Design and development of an on-line 1024 channel data acquisition and monitoring system for the CLRO correlators has been taken up, using a PC-AT as a basis.

Fabrication of first stage low-noise amplifiers and the cabling net-work for combining the antennas into groups is also in full progress.

#### Optical interferometry

During the past year a small group has been investigating the posibility of building a long base line optical interferometer. Before doing a feasibility study, it was proposed to gain first hand experience by setting up a Michelson interferometer in the campus. A pilot version of this with two 10 cm mirrors with a separation of 1.5 m has been set up. A suitable tracking system has been designed and testing is in progress. The design of an optical delay line with a resolution of 1 micron has been undertaken. With a view to selecting a suitable site for the full-scale interferometer, some 'seeing' measurements were done at Gauribidanur, Nandi Hills, Kavalur, etc., using a 15 cm telescope, a CCD camera, and frame grabbing techniques.

Advanced training in research is being offered to the following teachers from other organisations under the UGC Faculty Improvement Programme:

Name	Topic of Study
P.R. Maheshwara Murthy Govt. Science College Bangalore	Electric and Magnetic Field effects in Liquid Crystals
H.P. Padmini The National College Bangalore	Some physical studies on Liquid crystals
Sharanabasava M. Khened Laxmi Venkatesh Desai College, Raichur	Experimental studies of Phase Transitions in Ferroelectric Liquid Crystals
Ph. D	
Awarded	
Name	Topic of Study
K.S. Dwarakanath	A synthesis study of the radio sky at decameter wavelengths
S.N. Karbelkar	Statistical methods of phase reconstruction of speckle interferometry
R. Pratibha	Investigations on Liquid Crystalline Mixtures.
V.A. Raghunathan	Investigations on Some Field Induced Instabilities in Nematic Liquid Crystals
S. Sridhar	Some exact analytic models of time dependent collisionless stellar systems
Submitted	
C. Nagabhushan	Experimental Studies of Phase Dielectric Properties of Liquid Crystals
V.N. Raja	High Pressure and X-ray Studies of Liquid Crystals
S. Somasekhara	Experimental Studies of Phase Diagrams of Liquid Crystals

#### **Publications**

The research work done by the staff of the Institute has been published in a number of journals. A list of publications that have already appeared and those submitted and in press is given at Annexure-I (Page 19).

# Conferences/Seminars and Meetings

The staff of the Institute visited various institutions in India and abroad and attended conferences and presented papers. In addition, 74 lectures were given by them elsewhere.

#### Gandhi Memorial Lecture

The Gandhi Memorial Lecture for 1989 was given by Swami Ranganathananda on "The uniqueness of man in the Upanishads and modern science", on 2nd October, 1989.

# Colloquia

The scientists of the Institute and visiting scientists both from within and outside the country gave 22 colloquia at the Institute on different topics during the year.

### Journal Club Meetings

Fourteen meetings were held during the year. Preprints, as well recently published papers, dealing with topics of great current interest were reviewed in the meetings. A new 'Weekly Science Meeting' has been started during the year. In the 'Radio Astronomy Meetings' series each meeting is devoted to an informal presentation and discussion of a preselected topic; the 'Weekly Science Meetings' are without a fixed agenda and topics covering the work being carried on at the Institute, topics of current interest, etc., are discussed very informally.

# In-house Discussion Meeting

An in-house discussion meeting was held on 7th April 1990 on the current research activities of all groups in the Institute.

## Visiting Scientists

A number of scientists from institutions within the country and outside visited the Institute during the year. Their names are listed following those of the scientific and technical staff of the Institute given towards the end of the report.

# Library

Six hundred and twenty eight new books were added to the library during the year. The total book collection now is 17,100. The library presently subscribes to one hundred and forty six periodicals and has a collection of 20,762 bound volumes. Of the periodicals subscribed, 21 are received by airmail and IAU circulars are received via electronic mail.

# General

The following grants were received from the Department of Science and Technology during the year:

	(Rs. in lakhs)
PLAN-RECURRING	125.07
PLAN-NON RECURRING NON PLAN – RECURRING	186.92 52.00
	<del></del>
Total	363.99
-	<del></del>

#### **STAFF**

The scientific and technical staff of the Institute is given below. Additions during the year are marked with an asterisk.

### V. Radhakrishnan

K.R. Anantharamaiah

P.G. Ananthasubramanian

Antony Joseph

Arvind K. Shenai\*

T.K. Babu

R. Bhandari

D. Bhattacharya†

S. Chandrasekhar

K. Chandrasekhara

S. Chanthrasekharan

Chitra M. Gokhale

A.A. Deshpande

K.S. Dwarakanath

R. Ganesan

B.R. Iyer#

G. Jayakumar

N. Jayaprakash

P.A. Johnson

U.D. Kini

S. Krishna

S. Krishnan+

S. Krishnaprasad

V. Lakshminarayanan

Lashmy P. Usha

N.V. Madhusudana

M.O. Modgekar@@

Mohd. Ateegulla

R. Nandakumar

H. Narayanan\*

B.V. Nataraia

R. Nityananda

P. Rajasekhar\*

P.S. Ram Kumar

F.S. Kalli Kulliai

P.N. Ramachandra

C. Ramachandra Rao

# Joint Astronomy Programme

Somnath Bhardwaj

Nimesh Patel

B. Ramesh

T.K. Sridharan

Umesh Salian

S. Ramaseshan

P. Ramachandran\*

K. Ramesh Kumar

G.S. Ranganath

B.R. Ratna%

N. Ravi Sankar\*

D.K. Ravindra

T.S. Ravishankar

G. Rengarajan

T.N. Ruckmongathan@

B.K. Sadashiva

J. Samuel%%

G. Sarabagopalan

N.V.G. Sarma

P.S. Sasi Kumar

M. Seethalakshmi\*

M. Selavamani

Shaii P.

K. Shanthi\*

S. Shivkumar Nair

C.S. Shukre

K. Smiles Mascarenhas

P.S. Somasundaram

G. Srinivasan

K. Subramanya

M.R. Subramanyam

H. Subramonyam

K. Sukumaran

K.A. Suresh

V. Suresh Rao

N. Udaya Shankar

Umesh S. Nayak\*

C.V. Vishveshwara++

Vivek Dhawan\*

M. Vivekanand##

### Resignations

V.K. Anuradha

Javanthi Ramachandran

T. Ramachandran

R. Vijayalakshmi

# Consultant Physicians

# M.R. Baliga A.R. Pai

### **Post-Doctoral Fellows**

R. Pratibha
V.A. Raghunathan
V.N. Raja
Rani P. Rao
Ravi Kulkarni\*
Sai K. Iyer
Senthil Kumar\*
Subash Karbelkar

### Research Fellows

Archana Ghode Geetha G. Nair V. Rajeswari D.S. Shankar Rao G.B. Sivakumar P.B. Sunil Kumar L. Uma Devi K. Usha Veena Prasad

- On leave with the Centre for High Energy Astrophysics, University of Amsterdam, The Netherlands.
- # On deputation with the Institut des Hautes Etudes Scientifiques, France.
- + On deputation with the National Aeronautical Laboratory, Bangalore.
- @@ On deputation with the Australia Telescope National Facility, Sydney, Australia.
- % On leave with Georgetown University, Washintgon DC, USA.
- @ On leave with N.V. Philips, The Netherlands.
- %% On leave with Syracuse University, New York, USA.
- ++ On deputation as Director, Bangalore Planetarium, Bangalore.
- ## On leave with the U.S. Naval Observatory, Mt. Wilson, California, USA.

# LIST OF VISITORS

P.A.G. Scheuer Cavendish Laboratory Cambridge, U.K. March 3-April 7, 1989

May 23-June 17, 1990

L. Kogan

V. Vassili

S. Gordeev

Space Research Institute

The USSR Academy of Science

Moscow USSR

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